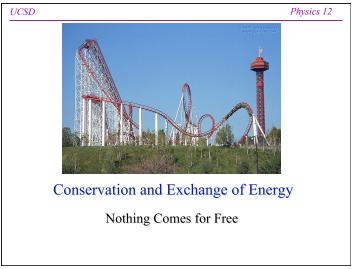
Energy Conservation Spring 2013



Energy is Conserved
Conservation of Energy is different from Energy Conservation, the latter being about using energy wisely
Conservation of Energy means energy is neither created nor destroyed. The amount of energy in the Universe is constant!!
Don't we create energy at a power plant?

Oh that this were true—no, we simply transform energy at our power plants

Doesn't the sun create energy?

Nope—it exchanges mass for energy

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Perpetual Motion

- Why won't the pendulum swing forever?
- It's hard to design a system free of energy paths
- The pendulum slows down by several mechanisms
 - Friction at the contact point: requires force to oppose;
 force acts through distance → work is done
 - Air resistance: must push through air with a force (through a distance) → work is done
 - Gets some air swirling: puts kinetic energy into air (not really fair to separate these last two)
- Perpetual motion means no loss of energy
 - solar system orbits come very close

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Lecture 5

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Some Energy Chains:

- A coffee mug with some gravitational potential energy is dropped
- potential energy turns into kinetic energy
- kinetic energy of the mug goes into:
 - ripping the mug apart (chemical: breaking bonds)
 - sending the pieces flying (kinetic)
 - into sound
 - into heating the floor and pieces through friction as the pieces slide to a stop
- In the end, the room is slightly warmer

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Bouncing Ball

- Superball has gravitational potential energy
- Drop the ball and this becomes kinetic energy
- Ball hits ground and compresses (force times distance), storing energy in the spring
- Ball releases this mechanically stored energy and it goes back into kinetic form (bounces up)
- Inefficiencies in "spring" end up heating the ball and the floor, and stirring the air a bit
- In the end, all is heat

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Gasoline Example

- Put gas in your car, containing 9 Cal/g
- Combust gas, turning 9 Cal/g into kinetic energy of explosion
- Transfer kinetic energy of gas to piston to crankshaft to drive shaft to wheel to car as a whole
- That which doesn't go into kinetic energy of the car goes into heating the engine block (and radiator water and surrounding air), and friction of transmission system (heat)
- Much of energy goes into stirring the air (ends up as heat)
- · Apply the brakes and convert kinetic energy into heat
- · It all ends up as waste heat, ultimately

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Why don't we get hotter and hotter

- If all these processes end up as heat, why aren't we continually getting hotter?
- If earth retained all its heat, we would get hotter
- All of earth's heat is *radiated* away

$$F = \sigma T^4$$

- If we dump more power, the temperature goes up, the radiated power increases dramatically
 - comes to equilibrium: power dumped = power radiated
 - stable against perturbation: T tracks power budget

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Rough numbers

- How much power does the earth radiate?
- $F = \sigma T^4$ for $T = 288^{\circ} \text{K} = 15^{\circ} \text{C}$ is 390 W/m²
- Summed over entire surface area $(4\pi R^2$, where R = 6,378,000 meters) is 2.0×10^{17} W
 - for comparison, U.S. production is 3×10^{12} W
- Solar radiation incident on earth is 1.8×10^{17} W
 - just solar luminosity of 3.9×10²⁶ W divided by geometrical fraction that points at earth
- Amazing coincidence of numbers! (or is it...)

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References & Assignments

Assignments

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- Read Chapter 2
- Homework #1 due Friday, April 12
- Homework #2 is now available:
 - go to Assignments page for link; I created all problems: none from book
 - start early on this one (toughest of quarter?)
- Quiz
 - nominally, deadline is Friday by midnight...
 - ...but still need to get TED revved up
 - monitor e-mail for announcement of availability
 - may extend deadline if delayed

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No Energy for Free

- No matter what, you can't create energy out of nothing: it has to come from somewhere
- We can *transform* energy from one form to another; we can *store* energy, we can *utilize* energy being conveyed from natural sources
- The net energy of the entire Universe is constant
- The best we can do is scrape up some useful crumbs

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Lecture 5